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RELATION BETWEEN EARTH TIDE OBSERVATIONS AND SOME OTHER DATA

por

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RELATION BETWEEN EARTH TIDE OBSERVATIONS AND SOME OTHER DATA

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1. Introduction.

Recently (Venedikov, 1988) it was proposed a new model for the study of the relation between Earth tide observations, say G and the air pressure data, say P . Here we shall give the results from the application of the model to a large record obtained in Pecny and to a shorter record in Sofia.

It is well known that the most important influence on the Earth tide observations is that of the ocean and sea tides. It is taken into account through computations made on a co-tidal map, now most often the map of Schwiderski. In such a way only the pure harmonic tidal oscillations of the water level are considered.

At a given coastal place it is certainly possible to have deviations in the water level. First, due to local particularities of the water basin the tides may be shifted in am-

plitude and phase compared to the tides predicted by the co-tidal map. Second, the level may be influenced by meteorological factors as air pressure and wind (meteorological tides). At the same time just the closest water masses have the strongest effect especially on the coastal Earth tide stations.

In principal in the model mentioned above we can use as P any phenomena which is expected to influence the Earth tide observations. In particular, for a coastal station, it could be interesting to use as P the water level. Thus we would obtain some direct estimates of the effect of the closest ocean or sea zones. It seems that this may be a helpful addition to the computations made on the co-tidal map.

We have realized such an experiment with the coastal gravity station Cueva de los Verdes (Lanzarote). The results are also given in the present paper.

2. The model used.

We shall briefly discuss the model with some precisising explanations.

Let \bar{S} is a vector representing a periodic component of P . We suppose that \bar{S} generates a component \bar{Q} with the same periodicity in the Earth tide record G . If $Q = |\bar{Q}|$ and $S = |\bar{S}|$ are the corresponding amplitudes and b is the angle (phase shift) between \bar{S} and \bar{Q} our hypothesis is

$$Q = B.S, \quad B = \text{const.} \quad \text{and} \quad b = \text{const.} \quad (1)$$

Let δ_s, χ_s

$$\xi_s = \delta_s \cos \alpha_s \quad \text{and} \quad \eta_s = -\delta_s \sin \alpha_s \quad (2)$$

are parameters of \bar{S} defined in a similar way as usual Earth tide parameters. This statement means that if we analyse P formally as if P are Earth tide data we shall obtain some estimates of the quantities (2).

Let $\bar{A}(\delta, \alpha)$ is an observed tidal component of G of the same type as \bar{S} , i.e. D , SD or TD . Observed means that \bar{A} incorporates the effect \bar{Q} of \bar{S} . And let $\bar{A}_0(\delta_0, \alpha_0)$ is the same component free from this effect, i.e.

$$\bar{A}_0 = \bar{A} - \bar{Q} \quad (3)$$

If we use the quantities

$$\begin{aligned} \xi &= \delta \cos \alpha, \quad \eta = -\delta \sin \alpha, \\ \xi_0 &= \delta_0 \cos \alpha_0, \quad \eta_0 = -\delta_0 \sin \alpha_0 \end{aligned} \quad (4)$$

the following simple relations come from (1) and (3)

$$\begin{aligned} \xi &= \xi_0 + B_1 \xi_s - B_2 \eta_s \\ \eta &= \eta_0 + B_1 \eta_s + B_2 \xi_s \end{aligned} \quad (5)$$

$$\text{where} \quad B_1 = B \cos b \quad \text{and} \quad B_2 = B \sin b \quad (6)$$

The quantities ξ and η as defined by the first equations of (4) are the unknowns used in the ordinary tidal analysis. If we want to take into the account the effect of P , respectively \bar{S} and \bar{Q} , we have to replace ξ and η by the new unknowns ξ_0 and η_0 and include two more unknowns the regression coefficients B_1 and B_2 .

However this cannot be directly done through the expressions (5) with constant ξ_s and η_s . The equations used in the analysis will become linearly dependent. This reflects simply the fact that in principle there are no means to separate from the tidal data a stable constant meteorological or any other kind of wave with a tidal frequency.

Nevertheless the expressions (5) can be used if we consider on the variation with time of \bar{S} , ξ_s and η_s .

In the first stage of the analysis we process a filtering of independent short intervals of the tidal record. Let T is the central epoch of a given filtered interval. Through the technics proposed by Venedikov (1981) and Xu Houze (1984) we can get $\xi_s(T)$ and $\eta_s(T)$ separately for D , SD and TD , through the filtering of the data P .

Let x_s and y_s are the mean values of $\xi_s(T)$ and $\eta_s(T)$ respectively and let

$$\Delta x_s(T) = \xi_s(T) - x_s, \quad \Delta y_s(T) = \eta_s(T) - y_s \quad (7)$$

Then the expressions (5) become

$$\begin{aligned}\tilde{\xi} &= \tilde{\xi}_0 + B_1 \Delta x_s(T) - B_2 \Delta y_s(T) \\ \tilde{\eta} &= \tilde{\eta}_0 + B_1 \Delta y_s(T) + B_2 \Delta x_s(T)\end{aligned}\quad (8)$$

where

$$\begin{aligned}\tilde{\xi}_0 &= \xi_0 + B_1 x_s - B_2 y_s \\ \tilde{\eta}_0 &= \eta_0 + B_1 y_s + B_2 x_s\end{aligned}\quad (9)$$

It is these equations (8) which can be used in the analysis. As a result directly from the processing we shall obtain estimates of the quantities (9) which are still charged by the total effect of \bar{S} , respectively of P . In the same time we shall get the estimates of the regression coefficients B_1 and B_2 . Later we can get the corrected numbers ξ_0 and η_0 from (9) i.e.

$$\begin{aligned}\xi_0 &= \tilde{\xi}_0 - (B_1 x_s - B_2 y_s) \\ \eta_0 &= \tilde{\eta}_0 - (B_1 y_s + B_2 x_s)\end{aligned}\quad (10)$$

3. Some results for the air pressure.

We have analysed formally as an Earth tide record a large air pressure series from Pecny prepared by Broz and Simon. It covers in total an interval of 15.9 years, 1.09.1970 - 30.07.1986 with the following more important interruptions:

4 11.1971 - 30.01.1974, 20.08.1975 - 20.04.1976 and 1.09.1981 - 25.01.1982. The results - amplitudes, phases and mean square errors (R.M.S.) are presented in Table 1. The original data are given in 0.1 mm mercury, while the amplitudes are expressed in mm mercury. The phases are related to the corresponding theoretical tidal phases.

Statistically significant and relatively important amplitudes are observed only at the waves S_1 , S_2 , S_3 and some close tides $-(P_1, S_1, K_1, PSI_1, PHI_1)$, (T_2, S_2, K_2) . The evident conclusion is that the main oscillations in the atmosphere are of meteorological and non-tidal origin. The only clearly observed tidal wave is M_2 with a small but very significant amplitude.

It is interesting that the relation $D/SD/TD$ for the standard deviation is similar to this relation for the analysis of Earth tide data.

These series, after a corresponding reduction of the length, were processed together with a large record of the gravimeter GS 15/228 at the same station Pecny provided by S.Holub, J.Broz and Zd.Simon. It covers a time interval of 8.4 years, 22.04.1976 - 25.09.1984. The neto amount of the data is 48 708 hourly ordinates. The results are presented in Table 2.

By the same gravimeter GS 15/228 observations were made in Sofia during 1 year. A record 12.12.1981 - 4.12.1982 was processed together with parallel air pressure data from the meteorological station in Sofia. The results are presented also in Table 2.

If we take into account the precision in Sofia there is a rather good coincidence in the coefficients B_1 and B for

D tides in the two stations. For SD the results in Sofia are not satisfactory.

The lower precision in the determinations of the coefficients in Sofia can be only partly explained by that the record is shorter.

4. Direct relation between Earth tide and sea tide observations.

The motivation for this experimental study was given in § 1 .

In the station Cueva de los Verdes (Lanzarote) Earth tide observations are made by the gravimeter LaCoste and Romberg No 434. The station is situated closely to a maregraphic station. We have used two parallel series from these stations with length 0.8 year - 11.07.1987 - 27.04.1988.

The maregraphic data were submitted to an analysis as gravity Earth tide data. The results are presented in Table 3 as they are outprinted by the computer. The amplitudes are expressed in cm.

The SD tides are more important and better determined than the D tides. Some of the D tides are statistically insignificant : NO1, J1 and OO1. The SD tides as well as M3 are all significant. It is interesting that there is perfect coincidence in phase between M2 and the tidal potential.

The results from the analysis of the gravity data are presented in Table 4. It is very remarkable that with the exception of M3 all factors are very small, especially for SD. It is possible that this is due to the coastal situation of the station and a strong particular influence of the sea.

References.

- Venedikov, A.P., 1981. Determination of the tidal parameters from short intervals in the analysis of Earth tide records. BIM, 85
- Venedikov, A.P., 1988. A model for the study of the effect of the air pressure on the Earth tide data. BIM, under press.
- Xu Houze, 1984. Harmonic analysis for short data using Venedikov's filters (in chinese). Crustal deformations and Earthquake, V.4, No 2, Wuhan.

In the same Table 4 are given the coefficients B , B_2 , B and the angle b which are determined by the common processing of both series of data. These coefficients are significant for D and SD .

It will be interesting to calculate the corrected parameters for the gravity $M2$ after the expressions (10).

As the phase shift α is small for $M2$ for both Earth and sea data from (10) we have approximately

$M2$:

$$\begin{aligned}\delta_o(\text{corrected}) &= \delta_o(\text{gravimetric}) - B_1 \cdot \delta(\text{maregraphic}) \\ &= 1.014 + 0.103 \times 1.130 = 1.130\end{aligned}$$

This is still somewhat low value of δ but, in our opinion, it is remarkable that the correction is in the right sense. It is easy to verify that the corrections after (10) of all more important tides are in the sense to rise considerably the values of δ .

Table 1

Station Pecny, Air pressure, tidal analysis.

wave	amplitude	R.M.S.	phase diff.	R.M.S.
D tides				
SIGQ	0.005	±0.007	24.269	± 78.843
2Q1	0.005	0.007	144.091	89.179
SIG1	0.01	0.007	256.746	39.766
Q1	0.017	0.007	-11.963	23.244
RO1	0.017	0.007	111.915	23.423
O1	0.011	0.007	98.762	34.709
TAU1	0.011	0.006	15.312	33.368
NO1	0.006	0.005	180.944	53.832
CHI1	0.012	0.006	112.561	30.704
PI1	0.003	0.007	76.278	121.361
P1	0.054	0.007	249.291	6.942
S1	0.066	0.010	41.231	9.124
K1	0.038	0.007	203.246	10.143
PSI1	0.010	0.007	-44.816	38.088
PHI1	0.014	0.006	103.809	25.059
TETA	0.005	0.006	190.929	72.125
J1	0.011	0.006	-11.079	30.267
SO1	0.020	0.006	236.171	18.109
OO1	0.014	0.005	98.155	21.583
NU1	0.008	0.005	267.313	35.263
SD tides				
EPS2	0.003	0.002	162.070	43.573
2N2	0.001	0.002	190.021	176.205
MU2	0.003	0.002	76.455	38.704
N2	0.001	0.002	11.905	98.759
NU2	0.005	0.002	64.106	21.820
M2	0.010	0.002	119.242	11.042
LAMB	0.002	0.002	107.308	61.656
L2	0.001	0.002	-30.095	63.038
T2	0.015	0.002	-48.854	7.237
S2	0.151	0.002	193.825	0.752
K2	0.020	0.002	184.632	5.159
ETA2	0.001	0.002	127.047	146.158
2K2	0.002	0.001	-1.850	29.157
TD tides				
M3	0.001	0.001	13.929	112.531
S3	0.010	0.001	209.238	4.716

Standard deviations D 1.67, SD 0.52 TD 0.22

Note: the amplitudes are in mm mercury.

Table 2

		B_1	R.M.S.	B_2	R.M.S.	B	b
Station Pecny	D	-0.416	± 0.007	0.024	± 0.007	0.417	176.74 ⁰
	SD	-0.235	0.023	0.032	0.024	0.237	172.86
Station Sofia	D	-0.471	0.047	0.093	0.045	0.481	168.78
	SD	-0.081	0.110	0.057	0.112	0.099	144.69

The unit of the coefficients B_1 , B_2 and B is $\mu\text{gal}/\text{mbar}$

Table 3

STATION DUNNEES MAREGRAPHYQUES ESPAGNE
29 9 0 N -13 26 0 E H OM

LEAST SQUARE ANALYSIS/VENEDIKOV/74, PROG. SV.-ICET-CUCARME, MELCHICR, VENEDIKOV
FILTERS ON 30 HRS, NR 102, 202, 301, ELIM. POWERS D 2 SD 1 TO 1
COMPONENTS S1 O1 S2 N2 M3
POTENTIAL CARTWRIGHT-TAYLER-EDDEN / COMPLETE DEVELOPMENT
COMPONENTS ORIENTATED TOWARDS THE ELLIPSOID (SKALSKY)
COMPUTER CENTRE UNIVERSITY COMPLUTENCE, MADRID
COMPUTER IBM 4381 PROCESSED ON 21.09.88

NORMALIZATION FACTOR -1.0000
INERTIAL CORRECTIONS INTRODUCED

TIME INTERVAL 0.8 YEARS 287.4 DAYS 5580 READINGS 8 BLCKS WEIGHTS
G*** 870711/870818 870903/871008 871011/871013 871110/871129
G*** 871202/880214 880218/880305 880307/880314 880316/880427

WAVE GROUP ARGUMENT	N WAVE	ESTIMATED		AMPLIT.		PHASE DIFF.	RESIDUALS		
		AMPL.	R.M.S.	FACTOR	R.M.S.		R.M.S.	AMPL.	PHASE
105.-139.	05 J1	1.527	0.201	0.3018	0.0397	120.522	7.484	6.77	168.8
143.-149.	26 J1	4.348	0.214	0.1046	0.0081	77.047	2.805	29.95	171.9
152.-158.	22 N01	0.252	0.190	0.1211	0.0513	-31.750	41.698	2.20	-176.5
161.-164.	13 P1	1.437	0.231	0.1169	0.0188	-27.185	5.401	12.92	-177.1
165.-168.	20 K1	5.335	0.216	0.1436	0.0058	-26.312	2.286	37.56	-176.4
172.-177.	22 J1	0.055	0.179	0.0263	0.0860	-85.978	187.254	2.41	-178.7
181.-183.	37 J01	0.066	0.115	0.0580	0.1014	202.267	97.621	1.38	-179.0
207.-209.	155 A2	57.353	0.834	1.0020	0.0146	-1.820	0.781	9.26	-168.7
207.-235.	41 2N2	1.778	0.032	1.2258	0.0362	33.340	1.690	1.00	101.4
243.-248.	24 N2	14.273	0.082	1.3024	0.0074	11.534	0.325	3.12	66.0
252.-258.	20 A2	64.674	0.090	1.1299	0.0016	0.000	0.079	1.73	180.0
262.-265.	14 L2	1.108	0.049	0.6846	0.0305	-21.290	2.567	0.94	-154.5
267.-274.	12 S2	23.854	0.078	0.8957	0.0029	-21.576	0.178	12.36	-134.8
275.-285.	38 K2	6.780	0.055	0.9357	0.0075	-17.564	0.463	2.82	-133.5
327.-332.	18 M3	0.274	0.029	0.2706	0.0294	237.592	5.854	1.14	-168.4

STANDARD DEVIATIONS 0 12.02 SD 4.19 ID 1.31
O1/K1 1.1441 1-G1/1-K1 0.9756 M2/G1 6.8517

Table 4

29 10 0 N -13 26 J E H 60M

CUEVA DE LOS VERDES (LANZAROTE) GRAVIMETRIC LCR. 431

LEAST SQUARE ANALYSIS/ VENEDIKOV/74, PROG. SV.-ICET-CUCARME, MELCHICR, VENEDIKOV
 FILTERS ON 30 HRS, NR 102, 202, 301, ELIM. POWERS D 2 SD 1 TC 1
 COMPONENTS S1 O1 S2 N2 M3
 POTENTIAL CARTWRIGHT-TAYLER-EDDEN / COMPLETE DEVELOPMENT
 COMPONENTS ORIENTATED TOWARDS THE ELLIPSOID (SKALSKY)
 COMPUTER CENTRE UNIVERSITY COMPLUTENCE, MADRID
 COMPUTER IBM 4381 PROCESSED ON 21.09.68

NORMALIZATION FACTOR -1.0000
 INERTIAL CORRECTIONS INTRODUCED

TIME INTERVAL 0.8 YEARS 287.4 DAYS 5580 READINGS E BLOCKS WEIGHTS
 G*** 870711/870313 870908/871003 871011/871013 871110/871129
 G*** 871202/880214 880218/880305 880307/880314 880316/880427

WAVE GROUP	ESTIMATED	AMPLIT.	PHASE	RESIDUALS
ARGUMENT	N WAVE	AMPL. R.M.S.	FACTOR R.M.S.	DIFF. R.M.S. AMPL. PHASE
105.-159.	65 J1	5.930	0.035	1.1721 0.0068 -C.205 C.275 C.07 -26.7
143.-149.	26 J1	27.953	0.057	1.1338 0.0021 -C.077 C.150 C.66-176.6
152.-158.	22 N01	2.363	0.027	1.1372 0.0132 C.156 0.676 0.05 171.8
161.-164.	13 P1	13.621	0.037	1.1077 0.0030 C.428 C.161 0.57 165.8
165.-168.	20 K1	41.296	0.033	1.1113 0.0010 C.581 C.053 1.07 156.8
172.-177.	22 J1	2.404	0.027	1.1572 0.0132 C.098 C.650 0.01 154.3
181.-183.	37 J01	1.251	0.018	1.0990 0.0157 1.564 C.610 0.06 148.8
207.-205.	155 A2	63.227	0.406	1.1047 0.0071 1.557 0.345 3.63 151.7

B1 = -0.0576 +- 0.0083 B2 = -0.0132 +- 0.0070 F = 0.0987 BETA = 187.69

207.-234.	41 A2	1.476	0.031	1.0175 0.0213 C.205 1.122 0.21 178.5
243.-248.	24 A2	11.122	0.077	1.0150 0.0070 2.397 C.365 1.67 163.9
252.-258.	20 A2	58.038	0.043	1.0140 0.0007 2.629 C.053 8.90 161.2
262.-265.	14 A2	1.605	0.027	0.9921 0.0164 3.654 C.655 C.25 159.6
267.-274.	12 S2	27.073	0.278	1.0167 0.0104 3.674 C.501 4.24 155.9
275.-285.	33 K2	7.330	0.059	1.0200 0.0095 3.556 0.465 1.13 156.0

B1 = -0.1032 +- 0.0213 B2 = -0.0413 +- 0.0213 F = 0.1112 BETA = 201.81

327.-362.	18 A3	1.069	0.016	1.0888 0.0159 1.570 0.827 0.10 17.3
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B1 = -0.0313 +- 0.0285 B2 = -0.0043 +- 0.0253 F = 0.0316 BETA = 188.67

STANDARD DEVIATIONS J 1.63 S1 1.30 TD 0.62
 O1/K1 1.0201 1-01/1-K1 1.1988 M2/C1 C.8571

WORKING GROUP ON HIGH PRECISION TIDAL DATA PROCESSING

BONN, OCTOBER 4 - 6, 1988

DETERMINATION OF SOME PARTICULAR WAVES IN THE EARTH TIDE DATA

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1. Introduction.

Recently Melchior & Ducarme (1986) and Melchior, Crossely, Dehant & Ducarme (1987) (hereinafter MA) have detected, in the observations made by the superconductive gravimeter in Brussels, some new particular waves (hereinafter SW). Their period range is between $13^h.166$ and $17^h.126$, a most significant wave being observed at $13^h.924$.

Evidently SW are of non-tidal origin. In the spectrum they are situated between the D and SD bands, considerably closer to SD.

Briefly the method for the determination of SW used by MA consists in the following. First the signal preliminary determined is subtracted. Then the residual curve thus obtained is submitted to power spectrum analysis.

Here we shall demonstrate another way to study SW which may be used in addition to the research done by MA. We shall present a few preliminary but more or less encouraging results.

We have an information about the periods (frequencies) of SW theoretically and/or experimentally (MA) determined. Then it becomes possible, even recomendable, to use the Method of the least squares to estimate the elements of SW. The main advantages of this method are (i) we can create and use a clear and flexible mathematical model (or models) and (ii) the results are accompanied by well defined estimates of their precision.

The physical sense of SW is not the object of our paper. We can only refer to the papers of MA, Gunn & Aldridge (1987), Aldridge & Lumb (1988) where one may find an interpretation and a corresponding bibliography. We shall simply note that SW have an extremely low power - amplitudes of the order of a few nanogals (10^{-12} of the gravity) - and that they may have variable and unstable amplitudes, phases and frequencies.

2. The analysis model.

The model of the useful tidal signal in the method of analysis (Venedikov, 1966, 1978) is

$$L(t) = \sum_{j=1}^m \left\{ \begin{array}{l} b_j \\ \sum_{i=a_j}^{b_j} H_i \cos (P_i + w_i t) \end{array} \right. + \gamma_j \cdot \sum_{i=a_j}^{b_j} H_i \sin (P_i + w_i t) \quad (1)$$

where

$L(t)$: tidal signal at time t ,

i : index or sequential number of a given tide,

H_1 : theoretical (undeformable Earth) amplitude,

P_1 : theoretical phase for an initial epoch for which $t = 0$,

w_1 : angular velocity,

a_j, b_j : the tides with indices

$$a_j \leq i \leq b_j \quad (2)$$

shape a tidal group j ($j = 1, 2, \dots, m$); it is supposed that all of them have one and the same tidal parameters δ_j and α_j , and

$$\xi_j = \delta_j \cos \alpha_j, \quad \gamma_j = -\delta_j \sin \alpha_j. \quad (3)$$

The quantities (3) are the unknowns which are determined by the analysis. From them we obtain δ_j and α_j .

Let us suppose that the tidal record incorporates an additional wave of type SW whose angular velocity is w . Analytically this can be represented by the addition to the right side of the expression (1) of a term like

$$\Delta L(t) = \xi \cdot H \cos (P + wt) + \gamma \cdot H \sin (P + wt). \quad (4)$$

This term is composed in such a way in order to be similar to the representation of the tidal waves in (1). Namely H and P are included as theoretical amplitude and phase but they can

be chosen absolutely arbitrary. The quantities ξ and η are unknowns about which we have in mind expressions similar to (3). If we obtain ξ and η through a processing of the data, then the observed amplitude h of the wave SW will be calculated after

$$\delta^2 = \xi^2 + \eta^2 \quad \text{and} \quad h = \delta \cdot H \quad (5)$$

In MA the periods are given, naturally, with a limited precision, for example $13^h.924$. Evidently one may expect that the last digit is not certain. The uncertainty arises if we suppose that the waves SW change with time. That is why we have made the following numerical speculation.

The single wave (4) is replaced by the sum of 5 waves with very close periods/angular velocities. If T is a given period, these 5 waves have the periods $T - 0.0010$, $T - 0.0005$, T , $T + 0.0005$ and $T + 0.0010$.

In the computer's program SV there are 504 tides from the Cartwright-Teyler-Eden development plus 3 "meteorological" waves (see Table II). All these waves are enumerated by the index $i = 1, 2, \dots, 507$. Thus the five waves have the indices $i = 508, \dots, 512$. The sum through which the representation (4) is replaced looks like

$$L(t) = \xi \sum_{i=508}^{512} H \cos(P + w_i t) + \eta \sum_{i=508}^{512} H \sin(P + w_i t) \quad (6)$$

where the arbitrary H and P remain constants for all 5 waves, while w_i are slightly different

For the period $T = 13.924^h$ we have choosen

Table I

i	periods h	w_1 o
508	13.9230	25.8565
509	13.9235	25.8556
510	13.9240	25.8546
511	13.9245	25.8537
512	13.9250	25.8528

3. Filters for the determination of SW.

The first stage of the analysis consists in the filtration of independent intervals (intervals without overlapping) from the record which is processed. The length of the intervals most often used is 48 hours but the program SV allows its arbitrary choice, for example 36 hours. In the present study for the determination of SW we have choosen two lengths: 36 and 42 hours. The first one was choosen as we have used it frequently for the determination of the usual D and SD tides, and the second one - as it is approximately multiple (3 times) of the period 13.924.

The filters were constructed through consecutive orthogonalization (Venedikov, 1978) The tables II represent the response (columns SW1) of the filters used. The first of the columns SW1 is for the even filter, the second one - for the odd filter. Only the most important components are included in the tables.

The filters eliminate an arbitrary linear drift independently for each one of the filtered intervals. Their response is just 1.00000 for the component $i = 508$, i.e. for the first of the 5 waves SW. The D tides are reduced through the perfect elimination of both O1 and S1. However, as the precision for the determination of SW must be a very high one, generally the reduction of D is not satisfactory (Q1, K1). This is taken into account in the analysis.

SW are very close to SD and it is not possible to be separated from them within intervals of such a length. Something more, the response of the filters for SD is higher than 1.00000. This is because SW are between D and SD. If we conceive SW, D and SD as n-dimensional vectors ($n = 36$ or 42) D and SD are approximately orthogonal and SW lays approximately in the hyperplane (D, SD) , between D and SD. The orthogonalization of SW ^{to} the D components O1 and S1, through which we obtain filters eliminating O1 and S1, turns SW closer to SD than to the initial position of SW.

4. Some results.

We have processed the series of data obtained by the superconducting gravimeter in Brussels. The data were kindly given at our disposal by prof. Melchior with the help of Mlle De Becker. They cover a 5.6 years interval: 21.04.1982 - 26.11.1987, a little bit larger than the series used by MA.

Table III represents an output from the program SV for SW with a mean period $T = 13.924^h$ (see Table I). It can be seen which tidal groups are included in the computation, i.e.

how is composed the model (1).

At first there is one group indicated by K_M which comprises all 205 D tides. It appeared to be necessary to include this group in the model because the separation of SW from D, as it was stated in the previous paragraph, is not a perfect one. This is proved by the fact that the δ factor obtained for this group is a significant one (1.245 ± 0.098). At the same time it is not necessary to use a more detailed separation because D as a whole is strongly suppressed by the filters.

Then there are 13 SD groups - the most detailed separation which is used by the tidal analysis. One can see that we have obtained a very good determination of these tides. After that there is a single group M_3 composed by all TD waves. The filters were not designed to eliminate TD.

At the end there is the group of five SW waves (Table I). The digits printed as the argument number of Doodson are of no meanings. Here one can notice that the intermediately value $\delta = 0.34$ is just 100 times the observed amplitude $h = 0.0034$ $\mu\text{gal} = 3.4$ nanogal. This is because the arbitrary value of the "theoretical" amplitude H in (4) and (6) was chosen $H = 100$.

The extremely high precision of near 1 nanogal is due to the accumulation in the model (6) of 5 waves with very close periods (Table 1). It can be accepted as realistic only if the spectrum is a continuous one over the band defined by the five waves (discussion at the time of the meeting of the Working group on high precision tidal data processing and particular remark by prof. Jentzsch). In fact if we want to attribute the result to a single wave, i.e. a single line in the spectrum we

have to multiply both the amplitude and its R.M.S. (mean square error) by a factor near 5 . For differences between the 5 frequencies as in Table I the factor is 4.8 .

In Table IV the amplitudes with their R.M.S. (after a multiplication by 4.8) for several periods are given. These are some of the periods obtained by MA. An exception is the first period in the Table IV . It has been chosen as the closest period which can be separated from 13.924 within a record of 5.6 years. It is an important fact that there is not a significant amplitude at this period.

If we use the Student's criterea the amplitude is a significant one only at 13.924. Thus we have only a partly confirmation of the results obtained by MA.

Nevertheless we are more inclined to consider this as encouraging. Our results are to be considered as preliminary ones and they were obtained mainly in order to develop the technology for this way of processing. We realize that there is still room for perfectioning the computation : (i) modification of the program in order to make easier and more operational the work when new particular waves are searched, (ii) better and more motivated choice of the filters including their lengths and eliminated components, (iii) perfectioning of the model , (iv) separation of the whole interval of the data on shorter intervals delimited by strong earthquakes as in MA, (v) introduction of a new tide-potential development, (vi) consideration of the air-pressure influence, (vii) experimentation on other long series of data etc.

- - - - -

We express our thanks to prof P. Melchior for initiating

our research on this interesting and important delicate problem as well as for giving at our disposal the high precision data from the superconducting gravimeter in Brussels.

References

- Aldridge, K.D. and Lumb, I.I., 1988. On the nature of long-period gravimetric data from the Earth's fluid outer core. Manuscript : 1 - 4.
- Gunn, S. and Aldridge, K.D., 1987. Inertial waves in a differentially rotating fluid. Symposium U2 - Instability within the Earth and core dynamics, XIX General Assembly, IUGG, Vancouver.
- Melchior, P. and Ducarme, B., 1986. Detection of inertial gravity oscillations in the Earth's core with a superconducting gravimeter at Brussels. Phys. Earth Planet. Int., 42 : 129-134.
- Melchior, P., Crossley, D.J., Dehant, V.P. and Ducarme, B., 1987. Have inertial waves been identified from the Earth's core? Symposium U2 - Instability within the Earth and core dynamics, IUGG XIX General Assembly, Vancouver.
- Venedikov, A.P., 1966. Une méthode pour l'analyse des marées terrestres à partir d'enregistrements de longueur arbitraire. Obs. Roy. Belg., Comm. No 250, S. Géoph. No 71 : 437 - 459.
- Venedikov, A.P., 1978. Analysis of Earth tide records (in Russian) Study of the Earth tides, KAPG, Bull. No 1, Budapest.

Table II.A

T M

745

TMS

[illegible]

Table II. A (continuation)

WAVES	F	M	SMI	SMI
307	257	500	412	27.9502
310	258	504	253	23.6093
311	259	505	251	20.0491
312	260	506	249	28.4808
313	261	507	247	28.5126
314	262	508	245	23.4655
315	263	509	243	23.5550
316	264	510	241	23.6113
317	265	511	239	23.7450
318	266	512	237	23.7119
319	267	513	235	23.7134
320	268	514	233	23.7022
321	269	515	231	23.6912
322	270	516	229	23.6802
323	271	517	227	23.6692
324	272	518	225	23.6582
325	273	519	223	23.6472
326	274	520	221	23.6362
327	275	521	219	23.6252
328	276	522	217	23.6142
329	277	523	215	23.6032
330	278	524	213	23.5922
331	279	525	211	23.5812
332	280	526	209	23.5702
333	281	527	207	23.5592
334	282	528	205	23.5482
335	283	529	203	23.5372
336	284	530	201	23.5262
337	285	531	199	23.5152
338	286	532	197	23.5042
339	287	533	195	23.4932
340	288	534	193	23.4822
341	289	535	191	23.4712
342	290	536	189	23.4602
343	291	537	187	23.4492
344	292	538	185	23.4382
345	293	539	183	23.4272
346	294	540	181	23.4162
347	295	541	179	23.4052
348	296	542	177	23.3942
349	297	543	175	23.3832
350	298	544	173	23.3722
351	299	545	171	23.3612
352	300	546	169	23.3502
353	301	547	167	23.3392
354	302	548	165	23.3282
355	303	549	163	23.3172
356	304	550	161	23.3062
357	305	551	159	23.2952
358	306	552	157	23.2842
359	307	553	155	23.2732
360	308	554	153	23.2622
361	309	555	151	23.2512
362	310	556	149	23.2402
363	311	557	147	23.2292
364	312	558	145	23.2182
365	313	559	143	23.2072
366	314	560	141	23.1962
367	315	561	139	23.1852
368	316	562	137	23.1742
369	317	563	135	23.1632
370	318	564	133	23.1522
371	319	565	131	23.1412
372	320	566	129	23.1302
373	321	567	127	23.1192
374	322	568	125	23.1082
375	323	569	123	23.0972
376	324	570	121	23.0862
377	325	571	119	23.0752
378	326	572	117	23.0642
379	327	573	115	23.0532
380	328	574	113	23.0422
381	329	575	111	23.0312
382	330	576	109	23.0202
383	331	577	107	23.0092
384	332	578	105	22.9982
385	333	579	103	22.9872
386	334	580	101	22.9762
387	335	581	99	22.9652
388	336	582	97	22.9542
389	337	583	95	22.9432
390	338	584	93	22.9322
391	339	585	91	22.9212
392	340	586	89	22.9102
393	341	587	87	22.8992
394	342	588	85	22.8882
395	343	589	83	22.8772
396	344	590	81	22.8662
397	345	591	79	22.8552
398	346	592	77	22.8442
399	347	593	75	22.8332
400	348	594	73	22.8222
401	349	595	71	22.8112
402	350	596	69	22.8002
403	351	597	67	22.7892
404	352	598	65	22.7782
405	353	599	63	22.7672
406	354	600	61	22.7562
407	355	601	59	22.7452
408	356	602	57	22.7342
409	357	603	55	22.7232
410	358	604	53	22.7122
411	359	605	51	22.7012
412	360	606	49	22.6902
413	361	607	47	22.6792
414	362	608	45	22.6682
415	363	609	43	22.6572
416	364	610	41	22.6462
417	365	611	39	22.6352
418	366	612	37	22.6242
419	367	613	35	22.6132
420	368	614	33	22.6022
421	369	615	31	22.5912
422	370	616	29	22.5802
423	371	617	27	22.5692
424	372	618	25	22.5582
425	373	619	23	22.5472
426	374	620	21	22.5362
427	375	621	19	22.5252
428	376	622	17	22.5142
429	377	623	15	22.5032
430	378	624	13	22.4922
431	379	625	11	22.4812
432	380	626	9	22.4702
433	381	627	7	22.4592
434	382	628	5	22.4482
435	383	629	3	22.4372
436	384	630	1	22.4262
437	385	631	0	22.4152
438	386	632	0	22.4042
439	387	633	0	22.3932
440	388	634	0	22.3822
441	389	635	0	22.3712
442	390	636	0	22.3602
443	391	637	0	22.3492
444	392	638	0	22.3382
445	393	639	0	22.3272
446	394	640	0	22.3162
447	395	641	0	22.3052
448	396	642	0	22.2942
449	397	643	0	22.2832
450	398	644	0	22.2722
451	399	645	0	22.2612
452	400	646	0	22.2502
453	401	647	0	22.2392
454	402	648	0	22.2282
455	403	649	0	22.2172
456	404	650	0	22.2062
457	405	651	0	22.1952
458	406	652	0	22.1842
459	407	653	0	22.1732
460	408	654	0	22.1622
461	409	655	0	22.1512
462	410	656	0	22.1402
463	411	657	0	22.1292
464	412	658	0	22.1182
465	413	659	0	22.1072
466	414	660	0	22.0962
467	415	661	0	22.0852
468	416	662	0	22.0742
469	417	663	0	22.0632
470	418	664	0	22.0522
471	419	665	0	22.0412
472	420	666	0	22.0302
473	421	667	0	22.0192
474	422	668	0	22.0082
475	423	669	0	21.9972
476	424	670	0	21.9862
477	425	671	0	21.9752
478	426	672	0	21.9642
479	427	673	0	21.9532
480	428	674	0	21.9422
481	429	675	0	21.9312
482	430	676	0	21.9202
483	431	677	0	21.9092
484	432	678	0	21.8982
485	433	679	0	21.8872
486	434	680	0	21.8762
487	435	681	0	21.8652
488	436	682	0	21.8542
489	437	683	0	21.8432
490	438	684	0	21.8322
491	439	685	0	21.8212
492	440	686	0	21.8102
493	441	687	0	21.7992
494	442	688	0	21.7882
495	443	689	0	21.7772
496	444	690	0	21.7662
497	445	691	0	21.7552
498	446	692	0	21.7442
499	447	693	0	21.7332
500	448	694	0	21.7222
501	449	695	0	21.7112
502	450	696	0	21.7002
503	451	697	0	21.6892
504	452	698	0	21.6782
505	453	699	0	21.6672
506	454	700	0	21.6562
507	455	701	0	21.6452
508	456	702	0	21.6342
509	457	703	0	21.6232
510	458	704	0	21.6122
511	459	705	0	21.6012
512	460	706	0	21.5902
513	461	707	0	21.5792
514	462	708	0	21.5682
515	463	709	0	21.5572
516	464	710	0	21.5462
517	465	711	0	21.5352
518	466	712	0	21.5242
519	467	713	0	21.5132
520	468	714	0	21.5022
521	469	715	0	21.4912
522	470	716	0	21.4802
523	471	717	0	21.4692
524	472	718	0	21.4582
525	473	719	0	21.4472
526	474	720	0	21.4362
527	475	721	0	21.4252
528	476	722	0	21.4142
529	477	723	0	21.4032
530	478	724	0	21.3922
531	479	725	0	21.3812
532	480	726	0	21.3702
533	481	727	0	21.3592
534	482	728	0	21.3482
535	483	729	0	21.3372
536	484	730	0	21.3262
537	485	731	0	21.3152
538	486	732	0	21.3042
539	487	733	0	21.2932
540	488	734	0	21.2822
541	489	735	0	21.2712
542	490	736	0	21.2602
543	491	737	0	21.2492
544	492	738	0	21.2382
545	493	739	0	21.2272
546	494	740	0	21.2162
547	495	741	0	21.2052
548	496	742	0	21.1942
549	497	743	0	21.1832
550	498	744	0	21.1722
551	499	745	0	21.1612
552	500	746	0	21.1502
553	501	747	0	21.1392
554	502	748	0	21.1282
555	503	749	0	21.1172
556	504	750	0	21.1062
557	505	751	0	21.0952
558	506			

T	WAVE	T _M	TMS
0	554	0	0
1	554	0	0
2	554	0	0
3	554	0	0
4	554	0	0
5	554	0	0
6	554	0	0
7	554	0	0
8	554	0	0
9	554	0	0
10	554	0	0
11	554	0	0
12	554	0	0
13	554	0	0
14	554	0	0
15	554	0	0
16	554	0	0
17	554	0	0
18	554	0	0
19	554	0	0
20	554	0	0
21	554	0	0
22	554	0	0
23	554	0	0
24	554	0	0
25	554	0	0
26	554	0	0
27	554	0	0
28	554	0	0
29	554	0	0
30	554	0	0
31	554	0	0
32	554	0	0
33	554	0	0
34	554	0	0
35	554	0	0
36	554	0	0
37	554	0	0
38	554	0	0
39	554	0	0
40	554	0	0
41	554	0	0
42	554	0	0
43	554	0	0
44	554	0	0
45	554	0	0
46	554	0	0
47	554	0	0
48	554	0	0
49	554	0	0
50	554	0	0
51	554	0	0
52	554	0	0
53	554	0	0
54	554	0	0
55	554	0	0
56	554	0	0
57	554	0	0
58	554	0	0
59	554	0	0
60	554	0	0
61	554	0	0
62	554	0	0
63	554	0	0
64	554	0	0
65	554	0	0
66	554	0	0
67	554	0	0
68	554	0	0
69	554	0	0
70	554	0	0
71	554	0	0
72	554	0	0
73	554	0	0
74	554	0	0
75	554	0	0
76	554	0	0
77	554	0	0
78	554	0	0
79	554	0	0
80	554	0	0
81	554	0	0
82	554	0	0
83	554	0	0
84	554	0	0
85	554	0	0
86	554	0	0
87	554	0	0
88	554	0	0
89	554	0	0
90	554	0	0
91	554	0	0
92	554	0	0
93	554	0	0
94	554	0	0
95	554	0	0
96	554	0	0
97	554	0	0
98	554	0	0
99	554	0	0
100	554	0	0

FILTERS LENGTH +2 NK 102,501,202,

Table II.B (continuation)

WAVES	W	SW1	SW1
301 227 505	27.0002	0.931855	1.108250
370 228 505	28.0093	0.936758	1.108472
371 239 505	28.0011	0.934313	1.108551
371 244 600	28.3367	0.934987	1.106655
385 245 605	28.4397	0.930179	1.106058
390 246 605	28.4808	0.936307	1.105371
397 247 605	28.5126	0.931826	1.104764
397 248 605	28.5556	0.931927	1.103955
402 253 605	28.9113	0.931906	1.053333
404 254 605	28.9453	0.943836	1.052053
409 255 605	28.9319	0.943553	1.090512
410 256 605	28.9041	0.914372	1.090420
411 256 605	29.0252	0.939659	1.088666
420 257 605	29.0002	0.934290	1.086831
420 258 605	29.4555	0.938159	1.065445
420 259 605	29.5265	0.917701	1.066647
438 261 605	29.3106	0.970619	1.054938
441 271 605	29.1179	0.972389	1.030800
444 272 605	29.3589	0.971305	1.027245
444 273 605	30.0000	0.971216	1.026313
444 274 605	30.0011	0.970258	1.019905
453 275 605	30.0321	0.970271	1.019120
454 276 605	30.0843	0.964493	1.015915
455 277 605	30.5557	0.972900	0.967332
461 285 605	30.6265	0.961700	0.958940
461 286 605	31.0180	0.954156	0.859398
461 295 605	31.1002	0.952845	0.888137
461 296 605	31.1745	0.941751	0.807377
463 327 605	41.9159	0.901524	0.333671
469 335 605	42.3374	0.904980	0.298552
469 336 605	42.9318	0.900111	0.250491
470 337 605	43.4162	0.902146	0.196044
470 338 605	44.0005	0.906509	0.137344
502 345 605	44.5142	0.903552	0.075335
502 346 605	45.0000	0.911750	0.020278
502 347 605	47.0002	0.903470	0.210839
507 348 605	48.0000	0.900644	0.000000
507 349 605	48.3355	0.900000	1.000000
509 350 605	48.3355	0.900000	1.000000
509 351 605	48.3355	0.900000	1.000000
509 352 605	48.3355	0.900000	1.000000
509 353 605	48.3355	0.900000	1.000000
509 354 605	48.3355	0.900000	1.000000
509 355 605	48.3355	0.900000	1.000000
509 356 605	48.3355	0.900000	1.000000
509 357 605	48.3355	0.900000	1.000000
509 358 605	48.3355	0.900000	1.000000
509 359 605	48.3355	0.900000	1.000000
509 360 605	48.3355	0.900000	1.000000
509 361 605	48.3355	0.900000	1.000000
509 362 605	48.3355	0.900000	1.000000
509 363 605	48.3355	0.900000	1.000000
509 364 605	48.3355	0.900000	1.000000
509 365 605	48.3355	0.900000	1.000000
509 366 605	48.3355	0.900000	1.000000
509 367 605	48.3355	0.900000	1.000000
509 368 605	48.3355	0.900000	1.000000
509 369 605	48.3355	0.900000	1.000000
509 370 605	48.3355	0.900000	1.000000
509 371 605	48.3355	0.900000	1.000000
509 372 605	48.3355	0.900000	1.000000
509 373 605	48.3355	0.900000	1.000000
509 374 605	48.3355	0.900000	1.000000
509 375 605	48.3355	0.900000	1.000000
509 376 605	48.3355	0.900000	1.000000
509 377 605	48.3355	0.900000	1.000000
509 378 605	48.3355	0.900000	1.000000
509 379 605	48.3355	0.900000	1.000000
509 380 605	48.3355	0.900000	1.000000
509 381 605	48.3355	0.900000	1.000000
509 382 605	48.3355	0.900000	1.000000
509 383 605	48.3355	0.900000	1.000000
509 384 605	48.3355	0.900000	1.000000
509 385 605	48.3355	0.900000	1.000000
509 386 605	48.3355	0.900000	1.000000
509 387 605	48.3355	0.900000	1.000000
509 388 605	48.3355	0.900000	1.000000
509 389 605	48.3355	0.900000	1.000000
509 390 605	48.3355	0.900000	1.000000
509 391 605	48.3355	0.900000	1.000000
509 392 605	48.3355	0.900000	1.000000
509 393 605	48.3355	0.900000	1.000000
509 394 605	48.3355	0.900000	1.000000
509 395 605	48.3355	0.900000	1.000000
509 396 605	48.3355	0.900000	1.000000
509 397 605	48.3355	0.900000	1.000000
509 398 605	48.3355	0.900000	1.000000
509 399 605	48.3355	0.900000	1.000000
509 400 605	48.3355	0.900000	1.000000
509 401 605	48.3355	0.900000	1.000000
509 402 605	48.3355	0.900000	1.000000
509 403 605	48.3355	0.900000	1.000000
509 404 605	48.3355	0.900000	1.000000
509 405 605	48.3355	0.900000	1.000000
509 406 605	48.3355	0.900000	1.000000
509 407 605	48.3355	0.900000	1.000000
509 408 605	48.3355	0.900000	1.000000
509 409 605	48.3355	0.900000	1.000000
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STATION BRUSSELS VERTICAL COMPONENT BELGIUM
 GRAVIMETER SUPERCONDUCTIVE, OBSERVATOIRE ROYAL DE BELGIQUE
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Table III

P.MELCHIOR

LEAST SQUARE ANALYSIS/VENEDIKOV/74, PROG.SV.-ICET-EUCARME, MELCHIOR, VENEDIKOV
 FILTERS ON 30 HRS, NR 102, 301, 202, ELIM. POWERS SW 1
 COMPONENTS S1 S2 S3 S4 S5 S6
 POTENTIAL CARLHANT-TAYLER-EDDEI / COMPLETE DEVELOPMENT
 COMPONENTS ORIENTATED TOWARDS THE ELLIPSOID (SKALSKY)
 COMPUTER CENTRE UNIVERSITY COMPLUTENCE, MADRID
 COMPUTER IOM 4381

PROCESSED ON 15.09.88

NORMALIZATION FACTOR 0.1000
 INERTIAL CORRECTIONS INTRODUCED

TIME INTERVAL 5.5 YEARS 2011.0 DAYS 48168 READINGS 3 BLOCKS NO WEIGHTS
 DATA 820421/820925 820902/861011 861115/871126

WAVE GROUP	ESTIMATED	AMPLITUDE	PHASE	RESIDUALS
ARGUMENT	WAVE AMPL. R.M.S.	FACTOR R.M.S.	DIFF. R.M.S.	AMPL. PHASE
105.-125. 205 K1	33.350 0.183	1.2452 0.0577	-0.097	4.467 4.61 -1.1
207.-229. 21 EP32	0.232 0.034	1.1350 0.0169	3.643	0.650 0.02 110.8
233.-255. 19 2N2	0.351 0.034	1.1120 0.0053	3.363	0.266 0.06 114.4
237.-259. 19 302	1.035 0.034	1.1932 0.0044	4.112	0.211 0.08 70.7
243.-245. 13 12	0.141 0.004	1.1730 0.0007	3.187	0.033 0.38 80.4
246.-248. 21 302	1.277 0.034	1.1700 0.0036	2.662	0.176 0.06 81.0
252.-258. 20 12	35.157 0.034	1.1011 0.0001	2.467	0.006 1.78 59.7
262.-264. 5 LA13	0.255 0.034	1.1437 0.0173	3.427	0.862 0.02 105.1
265.-267. 12 L2	0.119 0.034	1.1125 0.0033	2.505	0.164 0.05 110.5
271.-272. 2 12	0.134 0.034	1.2021 0.0040	1.128	0.216 0.04 29.6
273.-275. 4 32	15.902 0.034	1.2102 0.0003	1.132	0.013 0.77 25.6
274.-277. 12 K2	4.067 0.034	1.2070 0.0008	1.202	0.039 0.20 28.5
282.-285. 15 21A2	0.253 0.034	1.1834 0.0141	-0.142	0.678 0.01 -5.9
292.-295. 14 2K2	0.072 0.002	1.2348 0.0358	-1.943	1.593 0.01 -19.4
327.-375. 17 13	0.402 0.013	1.0738 0.0204	0.172	1.392 0.03 2.2
111.-111. 3 S31	0.0331 0.0012	0.31 0.12	256.77	22.96 0.01 -166.4

STANDARD DEVIATIONS SW 0.01

Table IV

FILTERS LENGTH	36		42	
	AMPLIT.	R.M.S.	AMPLIT.	R.M.S.
h				
13.918	7.6	± 5.8		
13.924	14.9	5.6	13.0	± 6.2
14.066	2.9	8.1	7.2	5.3
14.228	20.6	13.9	16.3	14.4
14.394	10.6	7.6		

The unit of the amplitudes is nanogal.

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DE LA UNIVERSIDAD COMPLUTENSE — MADRID

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